

Separation of parameterized and delayed sources: application to multispectral data

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Introduction

- Spectroscopy: data and applications
- Goal

2 Sparse-based Alternating Least Squares

3 Extension: Slow Delay Evolution Enforcement

4 Results and Comparisons

Galaxy kinematics

- The study of the intergalactic gas motion
- Data consist in multispectral images
- The peaks are shifted proportionally to the gas speed (Doppler effect)





Moment method

$$M_0 = \sum_{\lambda} X(\lambda), \quad M_1 = \frac{\sum_{\lambda} X(\lambda)\lambda}{M_0}, \quad M_2 = \sqrt{\frac{\sum_{\lambda} (\lambda - M_1)^2 X(\lambda)}{M_0}},$$



Goal



Goal



Challenges:

- Ill-posed inverse problem
- Similar and spectrally overlapping peaks
- Take into account the slow peak evolution

Source separation



I = 40 mixtures J = 4 sources

- Spectra *i* are assigned to mixtures
- Peaks j are assigned to sources
- Peak estimation and labeling are done simultaneously

Source separation



I = 600 mixtures J = 2 sources

- Spectra *i* are assigned to mixtures
- Peaks j are assigned to sources
- Peak estimation and labeling are done simultaneously

State of the art



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Mixing model



Estimate amplitudes A, delays L and shape parameters w that minimize the total residual error:

$$E(\boldsymbol{A}, \boldsymbol{L}, \boldsymbol{w}) = \left\| \boldsymbol{x}_i - \sum_{j=1}^J a_{ij} \boldsymbol{s}[\ell_{ij} \Delta; w_j] \right\|_2^2$$

Estimate amplitudes A, delays L and shape parameters w that minimize the total residual error:

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$$\boldsymbol{A} = \begin{bmatrix} a_{11} & \dots & a_{1J} \\ \vdots & \ddots & \vdots \\ a_{I1} & \dots & a_{IJ} \end{bmatrix} \quad \boldsymbol{L} = \begin{bmatrix} \ell_{11} & \dots & \ell_{1J} \\ \vdots & \ddots & \vdots \\ \ell_{I1} & \dots & \ell_{IJ} \end{bmatrix} \quad \boldsymbol{w} = \begin{bmatrix} w_1 & \dots & w_J \end{bmatrix}$$



ALS Scheme:

Until convergence:

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- **1** minimize $E(\boldsymbol{A}, \boldsymbol{L}, \boldsymbol{w})$ w.r.t. \boldsymbol{w}
 - Levenberg-Marquardt algorithm

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2 minimize E(A, L, w) w.r.t. A and L

Sparse approximation (OMP-like implementation)

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Sparse approximation (OMP-like implementation)

Separable problem w.r.t. to A and L:

$$\min_{\boldsymbol{A} \ge \boldsymbol{0}, \boldsymbol{L}} E(\boldsymbol{A}, \boldsymbol{L}, \boldsymbol{w}) \Leftrightarrow \forall i, \quad \min_{\boldsymbol{A}_{i:} \ge \boldsymbol{0}, \boldsymbol{L}_{i:}} \left\| \boldsymbol{x}_{i} - \sum_{j=1}^{J} a_{ij} \boldsymbol{s}[\ell_{ij} \Delta; w_{j}] \right\|_{2}^{2}$$

Sparse approximation



OMP-like implementation for mixture *i*:

Till all the sources are selected:

$$(\cdot)_{+} = \max(\cdot, 0)$$

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- G Remove the selected delayed source from the residual

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- Comp. time: 1.9 s
- Ambiguity in the separation of sources with very similar shapes

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Regularized criterion

Motivations:

- New discriminating factor to separate highly correlated sources
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$$F(\boldsymbol{A}, \boldsymbol{L}, \boldsymbol{w}) = \underbrace{E(\boldsymbol{A}, \boldsymbol{L}, \boldsymbol{w})}_{\text{Data-fit}} + \tau \underbrace{\Delta^2 \sum_{(i,i') \in \mathcal{G}} \sum_{j=1}^{J} (\ell_{ij} - \ell_{i'j})^2}_{\text{Delay regularization}}$$

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 $\mathcal{G} \colon$ set of cliques



Alternating Least Squares

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Alternating Least Squares

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Levenberg-Marquardt algorithm

minimize F(A, L, w) w.r.t. A and L
Joint sparse approximation
Neighbor sparse solutions



Neighbor sparse solutions



Neighbor sparse solutions



Joint NN-OMP

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• Comp. time: 1 s

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Method comparison

SNR (0 \rightarrow 30 dB), 1000 datasets: I = 30 mixtures, J = 3 sources, N = 200 samples



$$\mathsf{MSE} = \frac{1}{N \cdot I} \sum_{i=1}^{I} \left\| \boldsymbol{x}_{i} - \sum_{j=1}^{J} \widehat{a}_{ij} \boldsymbol{s}[\widehat{c}_{ij}; \widehat{w}_{ij}] \right\|_{2}^{2}$$

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$$\mathsf{error} = \frac{\|\boldsymbol{\Theta}^* - \widehat{\boldsymbol{\Theta}}\|_F^2}{\|\boldsymbol{\Theta}^*\|_F^2}$$

• Processing of multispectral data

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- Goals: estimate the peak characteristics (amplitudes, positions, shapes) and track their evolution
- Parameterized source separation framework to achieve both goals simultaneously

• Source separation using a sparse-based ALS scheme

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- Source separation using a sparse-based ALS scheme
- Slow delay evolution enforcement by introducing a regularization and a joint sparse approximation algorithm
- Application to real data (galaxy kinematics)
- The proposed methods are more adapted to the problem than literature delayed source separation approach and much faster than a Bayesian method

Publications

- H. Mortada, V. Mazet, C. Soussen, C. Collet, and L. Poisson. "Parameterized source separation for delayed spectroscopic signals". Signal Processing, vol. 158, pp. 48-60, may 2019.
- H. Mortada, V. Mazet, C. Soussen, and C. Collet. "Separation of delayed parameterized sources". In: Proceedings of the European Signal Processing Conference, EUSIPCO, pp. 1080-1084, 2017.
- H. Mortada, V. Mazet, C. Soussen, and C. Collet. "Spectroscopic decomposition of astronomical multispectral images using B-splines". In: Workshop on Hyperspectral Images and Signal Processing, WHISPERS, 2018.
- H. Mortada, V. Mazet, C. Soussen, and C. Collet. "Séparation de sources retardées et paramétriques". In: GRETSI, 2017.



Cost function



\boldsymbol{w} initialization

Joint OMP with slow delay evolution

end

ICM strategy

$$\widetilde{\boldsymbol{L}}_{:j} \leftarrow \operatorname*{argmax}_{\boldsymbol{L}_{:j}} \sum_{i=1}^{I} \left(\boldsymbol{r}_{i}^{T} \boldsymbol{s}[\ell_{ij}\Delta; \widehat{w}_{j}] \right)_{+}^{2} - \tau \Delta^{2} \sum_{(i,i') \in \mathcal{G}} (\ell_{ij} - \ell_{i'j})^{2}$$

ICM strategy

• Sequentially sweep
$$i \in \{1, \dots, I\}$$

• $\tilde{\ell}_{ij} \leftarrow \operatorname*{argmax}_{\ell} \left(\boldsymbol{r}_i^T \boldsymbol{s}[\ell\Delta; \hat{w}_j] \right)_+^2 - \tau \Delta^2 \sum_{(i,i') \in \mathcal{G}} \left(\ell - \tilde{\ell}_{i'j} \right)^2$
• end

B-spline initialization

B-spline iterative initialization

• For
$$k = 1 \nearrow J$$

Solve:
 $(\widehat{\Phi}_{:,1:k}, \widehat{\Sigma}_{:,1:k}, \widehat{\Omega}_{:,1:k}) \leftarrow \underset{\Phi_{:,1:k}, \Sigma_{:,1:k}, \Omega_{:,1:k}}{\operatorname{argmin}} L(\Phi, \Sigma, \Omega)$
using a SQP algorithm, initialized as follows:
 $\Phi_{:,1:k}^{0} = [\widehat{\Phi}_{:,1:k-1} \quad \phi_{k}^{0}] \text{ with } \phi_{k}^{0} \sim \mathcal{U}(0, +\infty)$
 $\Omega_{:,1:k}^{0} = [\widehat{\Omega}_{:,1:k-1} \quad \omega_{k}^{0}] \text{ with } \omega_{k}^{0} \sim \mathcal{U}(w_{\min}, w_{\max})$
 $\Sigma_{:,1:k}^{0} = [\widehat{\Sigma}_{:,1:k-1} \quad \sigma_{k}^{0}] \text{ with } \sigma_{k}^{0} = \lambda_{\max} \mathbf{1}_{M}$

 λ_{max} is the wavelength with the highest energy in the average mixture:

end

NN-OMP vs NN-FPG

• SNR: 5, 15, 25 dB, 200 datasets with I = 1, J = 3, N = 200• TP = Card($S^* \cap \hat{S}$), FP = Card(\hat{S}) – TP

SNR		NN-OMP	NN-FPG
	MSE	$47 \cdot 10^{-3}$	$44 \cdot 10^{-3}$
5 dB	Sparsity	4.3	14.1
	TP	0.5	0.5
	FP	3.8	13.6
	MSE	$4 \cdot 10^{-3}$	$4 \cdot 10^{-3}$
15 dB	Sparsity	5.2	12.2
	TP	1.3	0.9
	FP	3.9	11.3
	MSE	$0.5 \cdot 10^{-3}$	$0.4 \cdot 10^{-3}$
25 dB	Sparsity	6.1	12.7
	TP	2.0	1.2
	FP	4.0	11.5

Slow varying shape extension

$$\boldsymbol{x}_{i} = \sum_{j=1}^{J} a_{ij} \boldsymbol{s}[\ell_{ij}\Delta; p_{ij}\Theta] + \boldsymbol{n}_{i}$$

$$G(\boldsymbol{A}, \boldsymbol{L}, \boldsymbol{P}) = \underbrace{F(\boldsymbol{A}, \boldsymbol{L}, \boldsymbol{P})}_{F(\boldsymbol{A}, \boldsymbol{L}, \boldsymbol{P})} + \mu \Theta^{2} \sum_{(i,j) \in \mathcal{A}} \sum_{i=1}^{J} (p_{ij} - p_{i'j})^{2}$$

Data-fit and delay regularization

 $(i,i') \in \mathcal{G} j=1$

Shape regularization

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Slow varying shape extension

Joint OMP with slow delay evolution

• For
$$k = 1 \nearrow J$$

for $j \in \{1, ..., J\} \setminus \mathcal{J}$ do
 $(\tilde{L}_{:j}, \tilde{P}_{:j}) \leftarrow \underset{L_{:j}}{\operatorname{argmax}} \sum_{i=1}^{I} (r_i^T s[\ell_{ij}\Delta; p_{ij}\Theta])_+^2 - \tau \Delta^2 \sum_{(i,i') \in \mathcal{G}} (\ell_{ij} - \ell_{i'j})^2 - \mu \Theta^2 \sum_{j \notin \mathcal{J}} (p_{ij} - p_{i'j})^2 \text{ end}$
 $\hat{j} \leftarrow \underset{j \notin \mathcal{J}}{\operatorname{argmax}} \sum_{i=1}^{I} (r_i^T s[\tilde{\ell}_{ij}\Delta; \tilde{p}_{ij}\Theta])_+^2 - \tau \Delta^2 \sum_{(i,i') \in \mathcal{G}} (\tilde{\ell}_{ij} - \tilde{\ell}_{i'j})^2 - \mu \Theta^2 \sum_{(i,i') \in \mathcal{G}} (p_{ij} - p_{i'j})^2$
 $\mathcal{J} \leftarrow \mathcal{J} \cup \{\hat{j}\}$
 $\hat{J} \leftarrow \underset{i=1}{\mathcal{J}} (\mathbf{L}_{:\hat{j}}, \tilde{P}_{:\hat{j}}) \leftarrow (\tilde{L}_{:\hat{j}}, \tilde{P}_{:\hat{j}})$ s.t. $\{A_{:\mathcal{J}} \ge \mathbf{0}, A_{:\overline{\mathcal{J}}} = \mathbf{0}\}$
for $i = 1 \nearrow I$ do $r_i \leftarrow x_i - \sum_{j \in \mathcal{J}} \hat{a}_{ij} s[\hat{\ell}_{ij}\Delta; \hat{p}_{ij}\Theta]$ end

end
Slow varying shape extension

ICM strategy

$$\begin{split} &(\widetilde{\boldsymbol{L}}_{:j},\widetilde{\boldsymbol{P}}_{:j}) \leftarrow \\ & \underset{\boldsymbol{L}:j,\boldsymbol{P}:j}{\operatorname{argmax}} \sum_{i=1}^{I} \left(\boldsymbol{r}_{i}^{T} \boldsymbol{s}[\ell_{ij}\Delta;p_{ij}\Theta] \right)_{+}^{2} - \tau \Delta^{2} \sum_{(i,i')\in\mathcal{G}} (\ell_{ij} - \ell_{i'j})^{2} - \mu \Theta^{2} \sum_{(i,i')\in\mathcal{G}} (p_{ij} - p_{i'j})^{2} \\ & \bullet \text{ Sequentially sweep } i \in \{1,\ldots,I\} \\ & \bullet \left(\widetilde{\ell}_{ij},\widetilde{p}_{ij} \right) \leftarrow \\ & \underset{\ell,p}{\operatorname{argmax}} \left(\boldsymbol{r}_{i}^{T} \boldsymbol{s}[\ell\Delta;p\Theta] \right)_{+}^{2} - \tau \Delta^{2} \sum_{(i,i')\in\mathcal{G}} \left(\ell - \widetilde{\ell}_{i'j} \right)^{2} - \mu \Theta^{2} \sum_{(i,i')\in\mathcal{G}} (p - \widetilde{p}_{i'j})^{2} \\ & \bullet \text{ end} \end{split}$$

Galaxy NGC-4254



Galaxy NGC-4254



Galaxy NGC-4254



Solver comparison (B-spline method)

100 datasets, ${\rm SNR}=15$ dB, I=40 mixtures, J=3 sources and N=150 samples.

	trust-region-reflective	active-set	interior points	SQP
MSE	1.1×10^{-1}	1.1×10^{-1}	1.2×10^{-1}	1.0×10^{-1}
Time (s)	4.12	4.88	3.5	4.7